

# Gravitational Waves and Predictions of K-Theory for Excited Gravitons in The Bulk

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**Abstract:** The discovery of gravitational waves and its theoretical connection with quantum gravity leads to the idea of M-theory realization of topological solitons in the framework of Type I string theory. Using K-theory for SO(n) and Sp(n) Type I string theory it was possible to calculate the Ramon-Ramon charges of D-branesoliton objects for extra dimensional space, that signalize about the presence of gauge fields of excited graviton type in the bulk.

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## I. Introduction

The discovery of gravitational waves (GW) on 14 September 2015 by LIGO and Virgo Collaborations and the confirmation of this discovery by later measurements during 2016-2017 was not only a vivid confirmation of Einstein's general theory of relativity, but also emphasized the generality of the four types of interactions in high-energy physics.

General theory of relativity (GTR) of Albert Einstein was completed in 1915 and the article "The Foundation of the General Theory of Relativity" was published in 1916 year<sup>1</sup>, where he regarded the gravitational force not as the exterior one but as the "space-time" property. His geometric theory of gravitation was confirmed by numerous experiments, including the explanation of the anomalous perihelion advance of the planet Mercury and by the deflection of starlight by the Sun in May 29, 1919<sup>2</sup>. GTR predicted the existence of GW and the appearance of irremovable physical divergences when considering black holes and space-time singularities<sup>3</sup>. The existence of the particle-wave dualism predicted by Louis de Broglie for three types of interactions (strong, weak and electromagnetic) and the wave-particle opening for gravitational interaction is an intuitive confirmation of the possibility of combining of all four kinds of interactions at high energies.

In this article we will consider the base of GTR, derive from it the formula for GW and consider the perspectives of the metric variations for the discovery of Kaluza-Klein particles at the LHC. In the framework of type I string theory we will calculate Ramon-Ramon charges, for orthogonal and symplectic groups, SO(n); Sp(n), and show the existence of the solitonic objects of Kaluza-Klein type in the space of extra dimensions.

## II. General Theory Of Relativity

The feature of discovery of GW from the merging of black holes pair is the simultaneous confirmation of their existence: GW and black holes. Einstein's equivalence principle, or the universal equality of inertial and passive gravitational masses leads to the need for a transition from the Mikowski metric to the space with pseudo-Riemannian one. In this theory objects move along the geodesics and intervals in four-dimensional space-time

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

are given by the components of the metric tensor

$$g_{\mu\nu} = \begin{pmatrix} g_{00} & g_{01} & g_{02} & g_{03} \\ g_{10} & g_{11} & g_{12} & g_{13} \\ g_{20} & g_{21} & g_{22} & g_{23} \\ g_{30} & g_{31} & g_{32} & g_{33} \end{pmatrix}$$

Metric components determine the curvature of the space

$$R = g^{\mu\nu} R_{\mu\nu}$$

where  $R_{\mu\nu}$  is Ricci tensor related to Riemann curvature tensor  $R^{\alpha}_{\mu\alpha\nu}$  as

$$R_{\epsilon\beta} = R^{\rho}_{\alpha\rho\beta} = \partial_{\rho}\Gamma^{\rho}_{\beta\alpha} - \partial_{\beta}\Gamma^{\rho}_{\rho\alpha} + \Gamma^{\rho}_{\rho\lambda}\Gamma^{\lambda}_{\beta\alpha} - \Gamma^{\rho}_{\beta\lambda}\Gamma^{\lambda}_{\rho\alpha}$$

and  $\Gamma^{\alpha}_{\beta\gamma}$  are Christoffel symbols derived from the metric

$$\Gamma^{\alpha}_{\beta\gamma} = \frac{1}{2} \left( \frac{\partial g_{\alpha\beta}}{\partial x^{\gamma}} + \frac{\partial g_{\alpha\gamma}}{\partial x^{\beta}} - \frac{\partial g_{\beta\gamma}}{\partial x^{\alpha}} \right).$$

Thus, according to GTR a massive objects bend the space-time and move along this curved ways. Einstein wrote: "matter tells space how to bend; space tells matter how to move"<sup>4</sup>. Einstein's field equations are the following

$$R_{\mu\nu} - \frac{R}{2} g_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}.$$

The left side of it is the geometric properties of space-time ( $\Lambda$  is the cosmological constant) and the right part is expressed through the energy-momentum tensor,  $T_{\mu\nu}$ , of the matter which includes the density and flux of energy and momentum in space-time<sup>5</sup>.

### III. Gravitational Waves

GW was predicted by Einstein in 1916<sup>6</sup> from the GTR as gravitational radiation similar to electromagnetic radion. The essence of his work consisted in the addition to Kronecker delta the small metric tensor  $\gamma_{\mu\nu}$  :

$$g_{\mu\nu} = -\delta_{\mu\nu} + \gamma_{\mu\nu}$$

and in using an ansatz

$$\gamma'_{\mu\nu} = \alpha_{\mu\nu} f(x_1 + ix_4)$$

with  $\alpha_{\mu\nu}$  - constants and  $f$  - real functions of  $(x_1 + ix_4)$ . Let us consider how this addition to the metric leads to GW. As is known,  $g_{\mu\nu} = \delta_{\mu\nu} + h_{\mu\nu}$ , where  $\delta_{\mu\nu}$  is the Lorentz metric and  $h_{\mu\nu}$  is a first order quantity,

$h = h^{\alpha}_{\alpha} = \delta^{\sigma\lambda} h_{\sigma\lambda}$ . We'll rewrite the Einstein equation of GTR with this addition to the metric. As

$$R_{\mu\nu} = -\frac{1}{2} \delta^{\sigma\lambda} h_{\mu\nu,\sigma\lambda} - \frac{1}{2} (h_{,\mu\nu} - h^{\beta}_{\mu,\nu\beta} - h^{\beta}_{\nu,\mu\beta}) \tag{1}$$

then the equation

$$R_{\mu\nu} - \frac{R}{2} g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

transforms into the following expression

$$-\frac{1}{2} \delta^{\sigma\lambda} h_{\mu\nu,\sigma\lambda} + g_{\mu\nu} \frac{1}{4} \delta^{\sigma\lambda} h_{,\sigma\lambda} = \frac{8\pi G}{c^4} T_{\mu\nu} \tag{2}$$

with the additional condition that the second term in (1) equal to zero,

$$\left( h^{\nu}_{\mu} - \frac{1}{2} \delta^{\nu}_{\mu} h \right)_{,\nu} = 0. \tag{3}$$

Equations (2) and (3) can be rewritten into the equations

$$\square \varphi^{\nu}_{\mu} = -\frac{16\pi G}{c^4} T^{\nu}_{\mu} \tag{4}$$

and  $\varphi^{\nu}_{\mu,\nu} = 0$  with  $\varphi^{\nu}_{\mu} = h^{\nu}_{\mu} - \frac{1}{2} \delta^{\nu}_{\mu} h$  accordingly.

The question of the equation for particles with  $m = 0$  and spin = 2 was investigated by W. Pauli and M. Fierz<sup>7</sup>.

The relativistic description of such particles is connected with the equation  $\square \psi^{\nu}_{\mu} = 0$  with the additional conditions  $\psi^{\nu}_{\mu,\nu} = 0$ . This equation is the (4) one with the energy-momentum tensor equal to zero. The solution of this equation is

$$\varphi \sim \frac{1}{r} = P_0(\cos\gamma)\frac{1}{r_1} + P_1(\cos\gamma)\frac{r}{r_1^2} + P_2(\cos\gamma)\frac{r^2}{r_1^3} + \dots$$

$\frac{1}{R}$

As is known, the expression for quadrupole interaction with multipole moments  $q, p, Q_{ik}(n = \frac{1}{R})$  is the following

$$\varphi = \frac{q}{R} + p\frac{n}{R^2} + 3Q_{ik}\frac{n_i n_k}{R^3}.$$

Therefore, we have the solution for the potential of the external field, which is created by some remote system (in our case these are two merging black holes) and in which the quadrupole-quadrupole interaction is the most important one for uncharged systems. At the end of this section, it would be interesting to compare gravitational and electromagnetic interactions, presented in Table 1, from<sup>8</sup>.

**Table no 1** : Comparison of electromagnetic and gravitational fields.

|                      | Electromagnetism               | Gravity  |
|----------------------|--------------------------------|--|
| field                | $A_\mu$                        | $\varphi_\mu^v$                                |
| source               | $J^\mu$                        | $T_\mu^v$                                      |
| gauge transformation | $A^\mu + \partial_\mu \lambda$ | $h_\mu^v - \frac{1}{2}\delta_\mu^v h$          |
| Lorenz gauge         | $\partial^\mu A_\mu = 0$       | $\varphi_{\mu,\nu}^v = 0$                      |
| field equation       | $\square A_\mu = 4\pi J_\mu$   | $\varphi_\mu^v = -\frac{16\pi G}{c^4} T_\mu^v$ |

## I. Quantum gravity and topological invariants of D-branes

### 1. M-theory and AdS/CFT correspondence

For description of gravity with strong influence on (quantum) matter, are formulated quantum field theories in curved space-time<sup>9</sup>. The appearance of singularities at points with large curvature in quantum-mechanical description of space-time led to the need for theory of quantum gravity. Despite the fact that there is no suitable theory that unifies all types of interactions, there exist a number of candidates<sup>10</sup>. One of such attempts which is characterized by an additional property as six extra dimensions in addition to the four-dimensional space-time is string theory<sup>11</sup>. Later, supergravity theory<sup>12</sup>, became a part of eleven-dimensional M-theory - consistent theory of quantum gravity. In paper<sup>13</sup> was considered the connection of M-theory and string theory with one dimension compactified on orbifold. There was stressed, that  $E_8 \times E_8$  heterotic string is related to an eleven-dimensional theory on the orbifold  $R^{10} \times S^1/Z_2$ , the 11D M-theory, when  $E_8$  gauge fields are at each 10D fixed point of this orbifold. The action of bosonic part is the following<sup>14</sup>

$$S = S_{SG} + S_{YM},$$

where  $S_{SG}$  is the familiar 11-dimensional supergravity and  $S_{YM}$  - the  $E_8$  gauge fields of Yang-Mills theories on the orbifold planes<sup>13</sup>. The metric of such theory can be represented by formula

$$ds_{11}^2 = (1+b)\eta_{\mu\nu} dx^\mu dx^\nu + R^2(1+\hat{\gamma})(dx^{11})^2 + V^{1/3}(\Omega_{AB} + h_{AB})dx^A dx^B,$$

where  $b, \hat{\gamma}$  and  $h_{AB}$  depend on  $x^{11}$  and on coordinates of the space of extra dimensions - Calabi-Yau manifold.  $V$  and  $R$  are moduli for Calabi-Yau volume and orbifold radius. The compactification of this theory on the Calabi-Yau manifold leads to 5D  $N = 1$  supersymmetric model with one dimension compactified on  $S^1/Z_2$ , from<sup>15</sup>. According to AdS/CFT correspondence<sup>16</sup>, the space-time can have more than three noncompact spatial dimensions if we live on a four-dimensional domain wall which is embedded in the higher dimensions. The Standard Model (SM) physics is defined in four-dimensional subspace - D3-brane and gravity can propagate in bulk space-time. In AdS geometry the extra dimension is strongly curved, so AdS/CFT correspondence is the "holographic" projection of the AdS to the gauge theory physics. Thus, we'll compare the properties of the space of extra dimensions with supersymmetric Yang-Mills theory in usual space-time. For example, in the AdS3 geometry, the Randall-Sundrum (RS) partition function is obtained by integrating over the bulk metric with two patches  $R_1$  and  $R_2$  of AdS:

$$Z_{RS}[\gamma] = e^{-2S_1} \left( \int_{R_1 \square R_2} Dg e^{-S_{EH}[g] - S_{GH}[g]} \right)$$

(where  $S_{EH} = \frac{L}{16\pi} \int_{\Omega} d^3x \sqrt{g} \left( R + \frac{2}{l^2} \right)$  - the usual Einstein-Hilbert (EH) action with a negative cosmological constant  $\left( \Lambda = -\frac{1}{l^2} \right)$ ,  $S_{GH} = \frac{L_p}{8\pi} \int_{\partial\Omega} d^2x \sqrt{\gamma} K$  - Gibbons-Hawking (GH) boundary term, that depends on the bulk metric (K is the trace of the extrinsic curvature of the boundary,  $\gamma$  is the induced metric on the brane),  $S_1$  is the brane action without matter,  $S_1 = (T/2) \int d^2x \sqrt{\gamma}$ . According to<sup>16</sup>, this partition function is identified with the functional of Green's functions of the RS CFT:

$$Z_{RS}[\gamma] = e^{-2W_{CFT}[\gamma] + 2S_2[\gamma]}$$

So, the RS-like model in  $AdS_3$  is equivalent to a CFT coupled to gravity with action of matter on the brane,  $S_2 \propto \int d^d x \sqrt{\gamma} R$ .

## 2. K-group calculations

According to the D-brane approach in M-theory, the ends of open strings are restricted to D4-branes in the 11-dimensional space-time. The SM fields -harmonics of open string are localized on the D4-brane. The excited graviton states are closed string modes and propagate in the bulk. In Fig. 1 is presented the space-time in the Type I' string theory<sup>17</sup>.

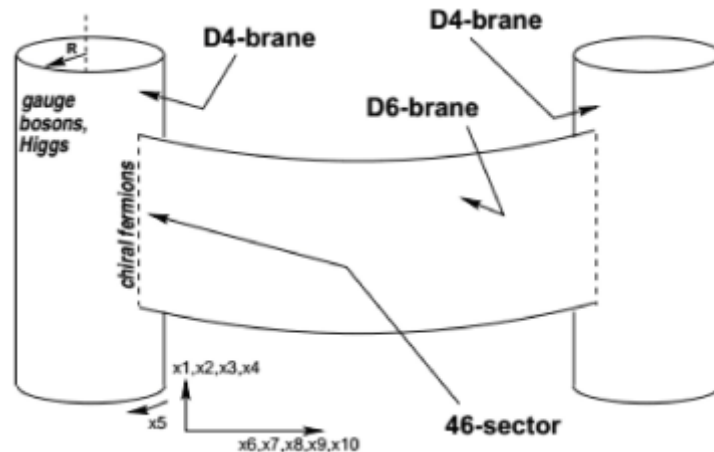


Figure no1: Schema of the space-time in the Type I' string theory.

The D4-brane corresponds to three space-time dimensions (usual space) plus one dimension compactified on the circle  $S^1$ . D6-brane corresponds to the rest of the dimensions, five of which are the spatial dimensions.

D-branes of type I are unoriented and their supersymmetric vacuum configuration has a richer spectrum of particles for calibration group. In paper<sup>17</sup> was shown that the unification scale can become lower to the TeV scale through the extra dimensions, which lead to gauge coupling unification at scales of TeV, presented in the theory of grand unification physics - string theory, Fig 2.

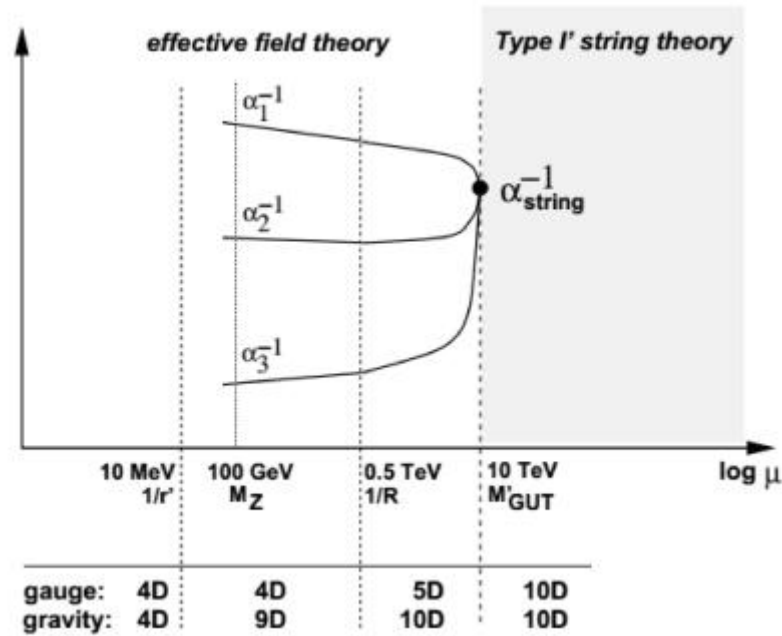


Figure no2: The evolution of the gauge couplings within a Type I' string theory.

The new unification scale is  $M_{GUT} \approx 10 \text{ TeV}$ , the gravitational coupling feels five new dimensions at  $(r')^{-1} \approx 10 \text{ MeV}$  and the gauge couplings feel a new dimension at  $R^{-1} \approx 0.5 \text{ TeV}$ , from<sup>17</sup>.

As is known, D-branes could be represented as fiber bundles<sup>18</sup>. We can compare the calculation of topological invariants of fiber bundles using the theory of categories<sup>19</sup>, with the properties of solitonic objects which are represented by such bundles<sup>20</sup>. As K-theory predicts the relations of the superstring theories and their D-brane spectra, for D-branes, considered as topological solitons, is used formalism of K-theory, which reproduces the spectra of BPS and non-BPS D-branes.

We will consider the construction of Grothendieck group for a classification of vector bundles over compact manifolds<sup>20</sup>. Let  $X=S^n$  be a compact manifold and let  $Vect(X)$  be the additive category of complex vector bundles over  $X$ . As  $I^k$  is the trivial bundle of rank  $k$  over  $X$ , i.e.  $I^k \sim X \times C^k$ , we can consider bundle  $E$  over  $X$  which is stably equivalent to bundle  $F$ , denoted by  $E \sim F$ , if  $E \oplus I^m \sim F \oplus I^m$ . The K-group of a compact manifold  $X$  can be defined as the Grothendieck group of the category  $K(X) \equiv K(Vect(X))$ :  $K(X) = Vect(X) \times Vect(X) / \sim$ , where  $Vect(X)$ - the set of isomorphism classes of complex vector bundles  $E \rightarrow X$  with some structure group. The equivalence relation in  $Vect(X) \times Vect(X)$  is according to  $(E; F) \sim (E'; F')$  such that  $E \oplus F' \oplus G \sim E' \oplus F \oplus G$  with vector bundle  $G \in Vect(X)$ .  $[E] = [F]$  in  $K(X)$  if  $E$  and  $F$  are stably equivalent. The elements of  $K(X)$  are virtual bundles, which represent functor from the category of compact topological spaces to the category of abelian groups,  $X \rightarrow K(X)$ .

In accordance with this known construction, we can consider the classification of Type I string theory for  $SO(n)$  and  $Sp(n)$  structure groups of fiberbundles with base  $X$ . As is shown in Fig.1, the schema of the space-time in the Type I' string theory is presented by two types of D-branes: D4-branes and D6-branes. The excited graviton or Kaluza-Klein states propagate in the bulk, as was highlighted above. So, the bulk or D6-branes are topological solitons with brane spectra, presented by Ramon-Ramon charges, related to Type I string theory spectra which can be calculated through K-group. This D6-branes are fiber bundles over  $X$  - noncompact space, which can be replaced by its one-point compactification,  $S^n$ ;  $X = S^n \times Z$ , through the addition of manifold  $Z$  at infinity. In the paper<sup>20</sup>, several types of string theories and their K-group classifications for the transverse space,  $S^n$  were considered. For Type I string theory K-groups were calculated for extra dimensional space associated with transverse space,  $S^n$ . Since we are working with five spatial dimensions, the transverse space for it is  $S^4$  and:

$$\tilde{K}(S^4) = Z \text{ for } SO(n) \text{ bundles for Type I spectrum of D5-branes; } \tilde{K}(S^4) = Z \text{ for } Sp(n) \text{ bundles for Type I spectrum of D5-branes.}$$

The nonzero value of K-group together with the relations of D-brane spectra and string spectra<sup>21</sup>, and according to AdS/CFT correspondence, we conclude about the existence of the excited states of gauge particles in the bulk.

#### IV. Conclusion

We have considered the string - Dp-branes -AdS correspondence, when Dp-branes are considered as p+1-dimensional hypersurfaces in space-time with the endpoints of open strings or as solitonic topological defects in space-time. According to<sup>22</sup>, D-branes as solitons are connected with fundamental string states and have provided the insights into the nature of space-time at very short distance scales. D-branes are also characterized by the determination of their mass by their charge and D-brane charges take values in K-group. As D-branes are extra dimensional objects, which are of different geometry, there are the AdS/CFT correspondence, that correlates the geometry of spacetime with a certain content of the fields of the Yang-Mills theory. Therefore, the nonzero elements of K-group signals about the presence of excited gauge fields, including the graviton in the bulk, for orthogonal, SO(n) and symplectic, Sp(n), gauge bundles, that represents D-brane solitonic objects.

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